

European Southern Observatory

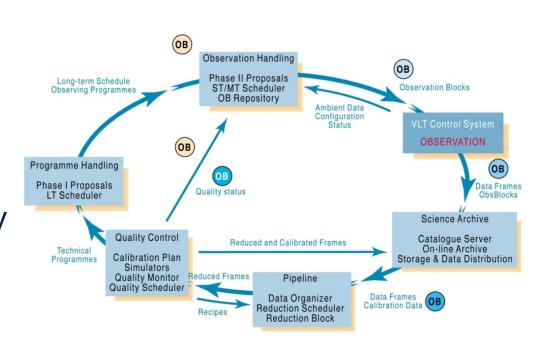
...on behalf of nearly 100 people in Germany and Chile who make it possible



Why an end-to-end process?

For reasons of observatory efficiency, ESO has implemented at the VLT and other current and future facilities (APEX, ALMA, VST, VISTA) Service mode observing at large scale (most or all the time).

In Service Mode, the main parts of an observing programme (writing a proposal/ planning and executing the observations/ reducing the data) are closely linked in an end-to-end model.





Writing the proposal

All you need to (and must!) know is in the Call for Proposals document, released twice per year: capabilities, policies, procedures...

Lots of tools to help you: Exposure Time Calculators, Archive, instrument technical webpages, information on calibration plans,...

And if this is not enough, help is offered by ESO: usd-help@eso.org





ESO Call for Proposals — P84
Proposal Deadline: 1 April 2009, 12:00 noon CEST



Lots of questions in the proposal form

(Instrument setup? Time on target? Execution conditions required? Time links?)...

..so that not only scientific merit and technical feasibility can be assessed:

also proper scheduling, taking into account weather expectation statistics, instrument availability, target distribution on the sky, etc.

The data in the proposal populate databases, thoroughly used afterwards

To efficiently schedule a programme, we (and you!) need to know as accurately as possible what you intend to do.

What happens next?



EUROPEAN SOUTHERN OBSERVATORY

Organisation Européenne pour des Recherches Astronomiques dans l'Hémisphère Austral Europäische Organisation für astronomische Forschung in der südlichen Hemisphäre

L.	Title								Category:	C-5
	Birth of a jet knot: Par-Lup3-4 seven years later									
	Abstra	ct / Tota	al Time Requ	ested						
	Total /	Amount (of Time:							
STREET STORES	object narrow that th typical more, velocit	Par-Lupi r-band im ie axis of i HH eject in the int y measur	4, discovered ages, only 1°2 the jet is almo fron velocity is ervening sever ements will al	ty us with from the ost contain is assumed. In years sin- llow us to e	th FORSI in 2 central object ed in the plan We thus exp ce the discovered determine its	t High: te of the pect a de ery of the total eje	that time of resolution of sky, leading early detects to knot, who ection veloci-	by the peculi- a knot in the je- pectroscopy ob- g us to infer a sable displacem ich together wi- ity. Given tha- nly one in this	t was clearly viservations have fast proper monent, of order of the available t Par-Lup3-4:	ristble in show stion if of 0°7 of le radio is one of
	motion			edy now, th		ons can		ful to constrain Sky Trans.	jet launch me	
Ų	A	84	FORS2	2h	mar	n	≤ 0.8°	THN	5	
			200							
			ghts/hours		Telescope	(s)		Amount of	time	
a)	alread	y awarded	ghts/hours to this projec complete this		Telescope -	(s)		Amount of	time	
a)	alread still re	y awarded	to this projec complete this		Telescope - -	n(s)		Amount of	time	
a) b)	alread still re	y awarded quired to	to this projec complete this		Telescope - -	(s)		Amount of -	time	
a) b)	alread still re Specia	y awarded quired to al remarks	to this projec complete this	project:		a(s)		Amount of	time	
a) b)	alread still re Specia	y awarded quired to al remarks	to this project complete this :: ::	project:		n(s)		Amount of	time	

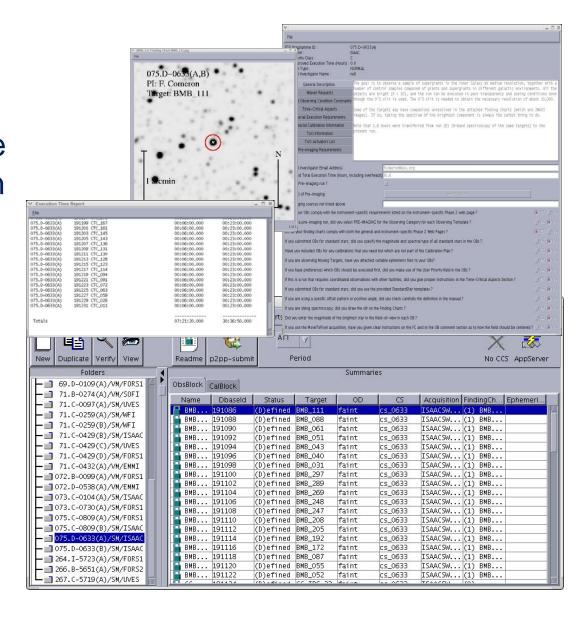
Is this proposal linked to a PhD thesis preparation? State role of PhD student



Now that you got observing time...

Phase 2 consists of the precise definition of the programme in the form of Observation Blocks that fully specify each composing observation

Done using the Phase 2
Proposal Preparation (P2PP)
tool





Phase 2 tools design

As far as possible, tools have been designed to ...



...allow users to do what they want to do

...be easy to use





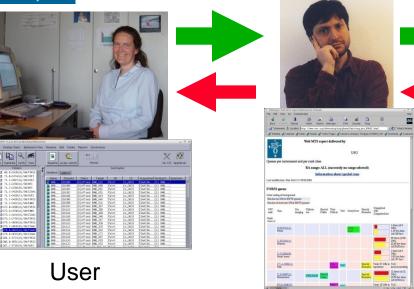
...help in avoiding mistakes

Phase 2 encourages you to think strategy, optimization, set up, execution conditions, priorities... in advance

A process closely assisted by ESO support staff: we are as interested as you in the efficient use of your telescope time!



How it works in Service Mode



Paranal Science Operations







User



User Support







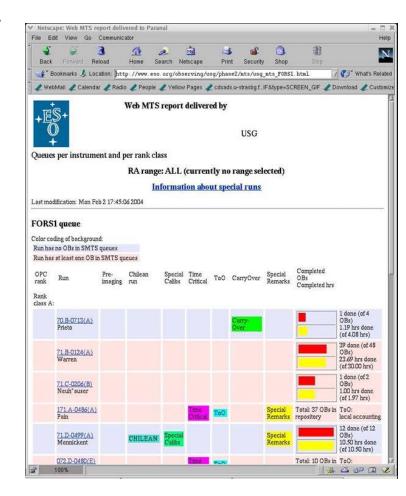
Quality control



In the queue

Once your programme has been validated and scheduled, execution will depend on

- Right conditions occurring
- Priority (A, B, C, Large Programmes,...)
- Pressure on the region of the sky
- Competition with special programmes (ToO, time-critical, pre-imaging...)
- Competition with programmes for other instruments
- Optimization of the night (trend to minimize calibration load)
- What has been done until now (preference to complete started programmes)



Progress and completion are hard to predict accurately...



Your data, finally...

All data undergo quality control, at the telescope and later off-line

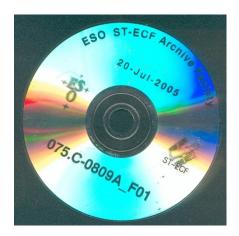
Important to monitor compliancy of the data with specifications, good performance of telescope and instrument, maintenance of the calibration database...

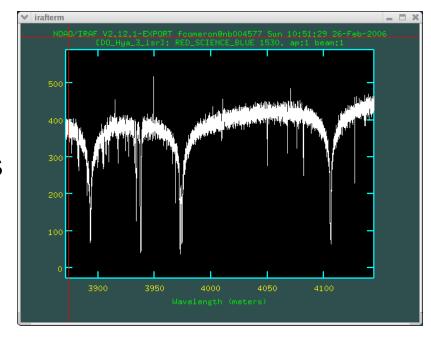
All the science data stream is processed by pipelines

Calibration plans are executed to ensure sufficient, uniform calibrations

All raw data, calibration data, logs, pipeline-processed data, and explanatory files are included in the data package









Proprietary data now accessible (to the PI!) as they reach the archive, within few hours of being taken (no pipeline-processed data though)

A FP7-funded pilot project, EVALSO, is currently being set up to allow much faster data transfer from Paranal, using fiber link connection from the observatory: data may be available in the future within seconds

Quick access



ESO Archive Query Results

ESO Archive Overview Help Page FAQ

Archive Facility HOME | ESO HOME

To request a single dataset press the button labeled Request Dataset next to the chosen dataset.

(You will be prompted for your ESO User Portal username and password. If you do not yet have an ESO User Portal account, please fill out the registration form.)

Datasets for which the proprietary period is over are highlighted in green and are publicly available.

Datasets that are still under the proprietary period are highlighted in red and can only be downloaded by the corresponding PI.

Define new query Status of Requests

MarkAll UnMarkAll

Request Marked Datasets Reset

<u>M</u>	More	<u>HDR</u>	Target Ra, Dec	Program ID	Instrument	Category	<u>Туре</u>	Mode	<u>Dataset ID</u>	Release
	®.	Header	16:09:03.03 - 39:05:00.8	081.C-0254 (B)	HAWKI	SCIENCE	OBJECT	IMAGE	HAWKI.2008- 07- 20T00:10:12.306	Oct 23 20
	•	Header	16:09:02.87 - 39:05:07.6	081.C-0254 (B)	HAWKI	SCIENCE	OBJECT	IMAGE	HAWKI 2008- 07- 20T00:11:01.954	Oct 23 20







Data legacy value

All data obtained with ESO telescopes are stored in the Archive

(but this is another story: Martino Romaniello's talk...)





At ESO, your observing programmes enter a system carefully designed to provide you with high quality data, optimizing the efficiency of the facility, and providing you support through all the phases along their lifetime ... and beyond